

Further Results of General Relativity

(GR) = a Summary of the most important Results.

קטגוריה 1, 13, 17

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$$dx^\alpha \rightarrow dx'^\alpha = \frac{\partial x'^\alpha}{\partial x^\beta} dx^\beta$$

$$A^\alpha \rightarrow A'^\alpha = \frac{\partial x'^\alpha}{\partial x^\beta} A^\beta$$

($\phi(x') = \phi(x)$) קטגוריה 1, 13, 17

$$\frac{\partial \phi}{\partial x^\alpha} \rightarrow \frac{\partial \phi}{\partial x'^\alpha} = \frac{\partial \phi}{\partial x^\beta} \frac{\partial x^\beta}{\partial x'^\alpha}$$

$$A_\alpha \rightarrow A'_\alpha = \frac{\partial x^\beta}{\partial x'^\alpha} A_\beta$$

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$$T^\alpha_\beta \rightarrow T'^\alpha_\beta = \frac{\partial x'^\alpha}{\partial x^\gamma} \frac{\partial x^\delta}{\partial x'^\beta} T^\gamma_\delta$$

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$$\frac{\partial}{\partial x^\alpha} V_\beta \rightarrow \frac{\partial}{\partial x'^\alpha} \left(\frac{\partial x^\delta}{\partial x'^\beta} V_\delta \right) \neq \frac{\partial x^\gamma}{\partial x'^\alpha} \frac{\partial x^\delta}{\partial x'^\beta} \frac{\partial V_\delta}{\partial x^\gamma}$$

$$\frac{\partial}{\partial x^\alpha} V^\beta \rightarrow \frac{\partial}{\partial x'^\alpha} \left(\frac{\partial x^{\beta'}}{\partial x^\gamma} V^\gamma \right) \neq \frac{\partial x^{\beta'}}{\partial x^\gamma} \frac{\partial x^\delta}{\partial x'^\alpha} \frac{\partial V^\gamma}{\partial x^\delta}$$

6.11.11 p 238 (2)

$$V^{\mu}_{;\lambda} \equiv \frac{\partial V^{\mu}}{\partial x^{\lambda}} + \Gamma^{\mu}_{\lambda\kappa} V^{\kappa}$$

$$V'^{\mu}_{;\lambda} = \frac{\partial x^{\mu'}}{\partial x^{\nu}} \frac{\partial x^{\rho}}{\partial x'^{\lambda}} V^{\nu}_{;\rho}$$

$$V_{\mu;\nu} \equiv \frac{\partial V_{\mu}}{\partial x^{\nu}} - \Gamma^{\lambda}_{\mu\nu} V_{\lambda}$$

$$V'_{\mu;\nu} = \frac{\partial x^{\rho}}{\partial x'^{\mu}} \frac{\partial x^{\sigma}}{\partial x'^{\nu}} V_{\rho;\sigma}$$

$$\Gamma^{\mu}_{\lambda\kappa} = \left\{ \begin{matrix} \mu \\ \lambda\kappa \end{matrix} \right\}$$

$$T^{\mu\sigma}_{;\lambda;\rho} = \frac{\partial T^{\mu\sigma}}{\partial x^{\rho}} + \Gamma^{\mu}_{\rho\nu} T^{\nu\sigma} + \Gamma^{\sigma}_{\rho\nu} T^{\mu\nu} - \Gamma^{\kappa}_{\lambda\rho} T^{\mu\sigma}_{;\kappa}$$

$$g_{\mu\nu;\lambda} = \frac{\partial g_{\mu\nu}}{\partial x^{\lambda}} - \Gamma^{\rho}_{\lambda\mu} g_{\rho\nu} - \Gamma^{\rho}_{\lambda\nu} g_{\rho\mu} = 0$$

1) α, β P'0 Wp

$$(\alpha A^{\mu}_{\nu} + \beta B^{\mu}_{\nu});_{\lambda} = \alpha A^{\mu}_{\nu;\lambda} + \beta B^{\mu}_{\nu;\lambda}$$

2) $(A^{\mu}_{\nu} B^{\lambda});_{\rho} = A^{\mu}_{\nu;\rho} B^{\lambda} + A^{\mu}_{\nu} B^{\lambda}_{;\rho}$

3) $g_{\mu\nu;\rho} = 0$; $g^{\mu\nu}_{;\rho} = 0$

$$(g^{\mu\nu} V_{\nu});_{\lambda} = g^{\mu\nu} V_{\nu;\lambda}$$

Riemann Christoffel 715)6

(3)

$$R^{\lambda}_{\mu\nu\kappa} \equiv \frac{\partial \Gamma^{\lambda}_{\mu\nu}}{\partial x^{\kappa}} - \frac{\partial \Gamma^{\lambda}_{\mu\kappa}}{\partial x^{\nu}} + \Gamma^{\eta}_{\mu\nu} \Gamma^{\lambda}_{\kappa\eta} - \Gamma^{\eta}_{\mu\kappa} \Gamma^{\lambda}_{\nu\eta}$$

$$R'^{\kappa}_{\rho\sigma\eta} = \frac{\partial x'^{\kappa}}{\partial x^{\lambda}} \frac{\partial x^{\mu}}{\partial x'^{\rho}} \frac{\partial x^{\nu}}{\partial x'^{\sigma}} \frac{\partial x^{\lambda}}{\partial x'^{\eta}} R^{\lambda}_{\mu\nu\kappa}$$

716) 715)6

$$T^{\mu\nu} : \quad \nabla_{\nu} T^{\mu\nu} = 0$$

Einstein 715)6

$$R_{\mu\nu} \equiv R^{\lambda}_{\mu\lambda\nu}, \quad R = g^{\mu\nu} R_{\mu\nu}$$

Ricci 715)6

Einstein 715)6

$$G_{\mu\nu} = R_{\mu\nu} - \frac{1}{2} g_{\mu\nu} R$$

$$G^{\mu\nu} = R^{\mu\nu} - \frac{1}{2} g^{\mu\nu} R$$

$$G^{\mu\nu}_{;\nu} \equiv 0, \quad \text{and } T^{\mu\nu}_{;\nu} = 0 \text{ is}$$

the covariant conservation of energy-momentum tensor.

Einstein. NICHE N

both sides $G^{\alpha\beta} = -8\pi G T^{\alpha\beta}$ have zero divergence $G^{\alpha\beta}_{;\beta} = T^{\alpha\beta}_{;\beta} = 0$

if $g_{00} \approx 1 + 2\phi$, $\delta_{ij} \approx -\delta_{ij}$

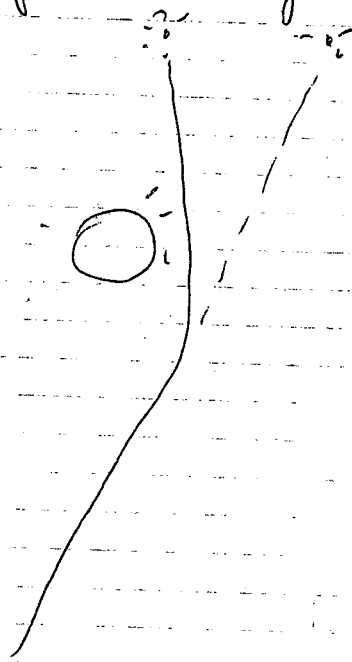
$$\begin{aligned} \Rightarrow \quad \nabla^2 g_{00} &= -8\pi G T_{00} \\ (\rho = T_{00}) \quad \nabla^2 \phi &= -4\pi G \rho \end{aligned}$$

The Schwarzschild solution of 4
empty Einstein's equations (exact solution)

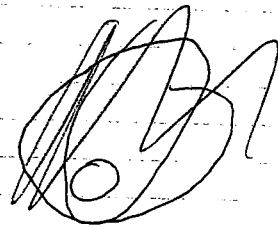
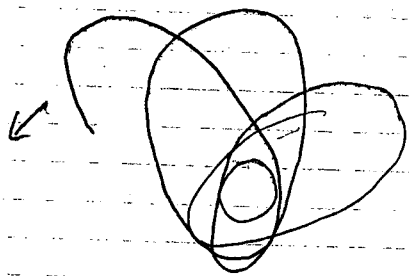
$$g_{00} = \left(1 - \frac{2GM}{r}\right), \quad g_{rr} = -\left(1 - \frac{2GM}{r}\right)^{-1}$$

$$g_{\theta\theta} = -r^2, \quad g_{\phi\phi} = -r^2 \sin^2\theta$$

- 1) Classical Tests: 1) Red shift (studied ✓)
- 2) Deflection of light



- 3) Precession Perihelion



Other Subjects to be studied.

(5)

(may be)
Cosmology & Expansion of the

Universe, Big Bang theory, Inflation theory
Observations of Accelerated Universe.

$$ds^2 = dt^2 - R(t)^2 (dx^2 + dy^2 + dz^2)$$

Geodesics "comoving observers" are for

example $x = x_0 = \delta^{12} p$

$$y = y_0$$

$$z = z_0$$

Distance between observers ("Galaxies").

$$\Delta l = R(t) |\Delta \vec{x}|$$

$$|\Delta \vec{x}| = \delta^{12} p$$

Gravitational Waves.

Action Principle & Symmetries

Other Theories of Gravity.